# THE ABERRATION OF LIGHT AND THE SPACE-TIME METRICS REDUCTION IN THE CASE OF ROTATION TO CANONICAL FORM

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Abstract. This paper presents the reduction to canonical form of the space-time metrics corresponding to a rotation of angular speed w round the axis Oz. An interpretation of this is also provided.

MSC 2000: 83C20

Keywords: space-time metrics, reduction to canonical form, aberration of light

### 1. Notations

In generalized spherical coordinates:  

$$x^0 = t \in \mathbb{R}, \quad x^1 = r \in [0, \infty), \quad x^2 = \theta \in [0, \pi], \quad x^3 = \phi \in \mathbb{R}$$
 (1)

in geometrized units [1]

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$$\begin{cases} t = ct_{ph}, & r = r_{ph}, & \theta = \theta_{ph}, & \phi = \phi_{ph}, \\ m = \frac{G_{ph}}{c^2} M_{ph}, Q \frac{G^{\frac{1}{2}}}{c^2} Q_{ph}, \omega = \frac{d\phi}{dt} = \frac{1}{c} \omega_{ph}, \\ v = \frac{1}{c} V_{ph}, a = \frac{1}{c^2} a_{ph} \end{cases}$$
 (2) 
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the expression of metrics for the space-time next to a spherical body is under the form

$$ds^{2} = g_{00}dx^{0^{2}} + g_{11}dx^{1^{2}} + g_{22}dx^{2^{2}} + g_{33}dx^{3^{2}}$$
(4)

The index "ph" signifies the fact that the respective quantities are measured in physical units,  $m, Q, \omega, v$ , a denote respectively the mass, electrical charge, anguler speed, linear velocity and the acceleration.

With the Reissner- Nordström metrics, the components of the fundamental metric tensor are: ) values of the strain and ground a selection of the

$$g_{00} = 1 - \frac{2m}{r} + \frac{Q^2}{r^2} > 0;$$
  $g_{11} = -\frac{1}{g_{00}} < 0;$   $g_{22} = -r^2 < 0;$   $g_{33} = -r^2 \sin^2 \theta < 0$  (5)

2. Determination of metrics QUA THOLE WOLLD ASSESSED SHIT Supposing a point  $A(r, \theta, \phi)$  rotates round the axis Oz at an angular speed

$$\omega = \frac{d\phi}{dt}$$

instead of the  $\phi$  coordinate we will have take  $\phi + \omega t$ 

For point A the following metrics will have to be taken:

$$ds^{2} = (g_{00} + \omega^{2}g_{33})(dt)^{2} + g_{11}(dr)^{2} + g_{22}(d\theta)^{2} + g_{33}(d\phi)^{2} + 2\omega g_{33}d\phi dt (6)$$

Noting with

$$v_0 = dt, v_1 = dr; v_2 = rd\theta; v_3 = r\sin\theta d\phi \tag{7}$$

the components of vector

$$\mathbf{v} = (v_0, v_1, v_2, v_3)$$
 (8)

from the vector space R4 metrics (6) becomes the real quadric form  $ds^2: R^4 \rightarrow R$ 

$$de^2 = a v^2 + a v^2 + a v^2 = -2 \cdot a$$

 $ds^2 = a_{00}v_0^2 + a_{11}v_1^2 + a_{22}v_2^2 + a_{33}v_3^2 + 2a_{03}v_0v_3$ (9)

where

here
$$\begin{cases}
a_{00} = 1 - \frac{2m}{r} + \frac{Q^2}{r^2} - \omega^2 r^2 \sin^2 \theta \ge 0, \\
a_{11} = -\left(1 - \frac{2m}{r} + \frac{Q^2}{r^2}\right)^{-1} \le 0, \\
a_{22} = a_{33} = -1 < 0, \quad a_{03} = -\omega r \sin \theta \le 0.
\end{cases} \tag{10}$$

The quadratic form matrix has the form

$$A = \begin{bmatrix} a_{00} & 0 & 0 & a_{03} \\ 0 & a_{11} & 0 & 0 \\ 0 & 0 & a_{22} & 0 \\ a_{03} & 0 & 0 & a_{13} \end{bmatrix}$$
(11)

# 3. Reduction to canonical form

The reduction of the quadratic form (9) to the canonical form is done through the method of eigenvalues and eigenvectors, as in this way we can determine an orthonormed basis corresponding to an isometry (rotation in R4).

From the characteristic equation

$$p(\lambda) = \left[\lambda^2 - (a_{00} - 1)\lambda - (a_{00} + a_{03}^2)\right](a_{11} - \lambda)(-1 - \lambda) \tag{12}$$

we get the eigenvalues:

$$\lambda_0 = \frac{a_{00} - 1 + \sqrt{\Delta}}{2} > 0; \lambda_1 = a_{11}; \lambda_2 = -1; \lambda_3 = \frac{a_{00} - 1 - \sqrt{\Delta}}{2} < 0$$
 (13)

where

$$\Delta = (a_{00} - 1)^2 + 4(a_{00} + a_{03}^2) > 0$$
 (14)

Metrics (9) aquires the form

$$ds^{2} = \lambda_{0}v_{0}^{2} + \lambda_{1}v_{1}^{2} + \lambda_{2}v_{2}^{3} + \lambda_{3}v_{3}^{2}$$
(15)

$$\dot{v_0} = dt', \dot{v_1} = dr', \dot{v_2} = r'd\theta', \dot{v_3} = r'\sin\theta'd\phi'$$
 (16)

As  $\lambda_1 = a_{11}$  and  $\lambda_2 = -1$  it results that

$$v_1' = v_1$$
, and  $v_2' = v_2$  (17)

Norming the eigenvectors that have been obtained from the characteristic system

$$\begin{cases} (a_{00} - \lambda)x_0 + a_{03}x_3 = 0\\ (a_{11} - \lambda)x_1 = 0\\ (-1 - \lambda)x_2 = 0\\ a_{03} - x_0 + (-1 - \lambda)x_3 = 0 \end{cases}$$
(18)

we get the vectors of the new basis

$$B' = \{e_0, e_1, e_2, e_3\},\tag{19}$$

$$\begin{cases} e_0' = \left(\frac{\alpha}{\sqrt{1+\alpha}}, 0, 0, \frac{-1}{\sqrt{1+\alpha^2}}\right) \\ e_1' = (0,1,0,0) \\ e_2' = (0,0,1,0) \\ e_3' = \left(\frac{+1}{\sqrt{1+\alpha^2}}, 0, 0, \frac{\alpha}{\sqrt{1+\alpha^2}}\right) \end{cases}$$
(20)

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$$\alpha = -\frac{a_{00} + 1 + \sqrt{\Delta}}{2a_{03}} > 0 \tag{21}$$

The matrix of transition from the canonical basis

$$B = \{e_0 = (1, 0, 0, 0), e_1 = (0, 1, 0, 0, ), e_2 = (0, 0, 1, 0), e_3 = (0, 0, 0, 1)\} (22)$$

to basis B' is

$$C = \begin{bmatrix} \frac{\alpha}{\sqrt{1 + \alpha^2}} & 0 & 0 & \frac{1}{\sqrt{1 + \alpha^2}} \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ -\frac{1}{\sqrt{1 + \alpha^2}} & 0 & 0 & \frac{\alpha}{\sqrt{1 + \alpha^2}} \end{bmatrix}$$
(23)

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(24)

we have a rotation in the space R<sup>4</sup> with the axes oriented e<sub>1</sub> and e<sub>2</sub> remaining unchanged. The rotation angle, γ, is determined from the formulae:

$$\cos \gamma = \frac{\alpha}{\sqrt{1 + \alpha^2}}; \sin \gamma = \frac{1}{\sqrt{1 + \alpha^2}} \text{ or } tg\gamma = \frac{1}{\alpha}$$
 (25)

Inserting \alpha from (21) and (14) in (25) we obtain

 $\det C = 1$ 

$$tg\gamma = \frac{\sqrt{(a_{00} + 1)^2 + 4a_{03}^2 - (a_{00} + 1)}}{-2a_{03}}$$
 (26)

namely

$$w = \omega r \sin \theta = \frac{w_{ph}}{c}$$
 (28)

is the velocity of point A in generalized units namely the aberration constant in nonrelativistic case. Taking m = 0 and Q = 0, we get

$$tg\gamma = \frac{\sqrt{w^4 + 4 + w^2 - 2}}{2w} = \frac{w}{2} \left( 1 + \frac{1}{4}w^2 - \frac{1}{64}w^6 + \frac{1}{516}w^{10} \dots \right) \approx \frac{w}{2}$$
 (29)

where in the first approximation in the half aberration constant

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