

**THE DOPPER EFFECT ASSOCIATED WITH THE SPACE-TIME  
METRICS REDUCTION IN THE RADIAL MOVE CASE CANONICAL  
FORM**

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**Abstract.** This paper presents the reduction to canonical form of the space-time metrics corresponding to a radial move at the speed  $w$ . A physics interpretation is also provided.

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**1. Notations**

In generalized and geometrized spherical coordinates:

$$x^0 = t, \quad x^1 = r, \quad x^2 = \theta, \quad x^3 = \phi \quad (1)$$

with the specifications and notations from the paper "The aberration of light and space-time metrics reduction in the case of rotation to canonical form" of the author of this publication the metrics of the quadridimensional space next to a spherical body has the form

$$ds^2 = g_{00}dx^{0^2} + g_{11}dx^{1^2} + g_{22}dx^{2^2} + g_{33}dx^{3^2} \quad (2)$$

Supposing a point  $A(r, \theta, \phi)$  next to the body started in the radial direction at the speed  $w$ . In metrics (2) we change  $dx^1$  for  $dx^1 + wdx^0$ . We obtain

$$ds^2 = (g_{00} + g_{11}w^2)dx^{0^2} + g_{11}dx^{1^2} + g_{22}dx^{2^2} + g_{33}dx^{3^2} + 2wg_{11}dx^0 dx^1 \quad (3)$$

Noting with

$$v_0 = dx^0, \quad v_1 = dx^1, \quad v_2 = x^1 dx^2, \quad (4)$$

$$v_3 = x^1 \sin x^2 dx^3$$

metrics (3) will be written as the quadratic form of  $R^4$ :

$$ds^2 = a_{00}v_0^2 + a_{11}v_1^2 + a_{22}v_2^2 + a_{33}v_3^2 + 2a_{01}v_0 v_1 \quad (5)$$

where

$$\begin{aligned} a_{00} &= g_{00} + g_{11}w^2; & a_{01} &= g_{11}; \\ a_{22} &= a_{33} = -1; & a_{01} &= wg_{11}. \end{aligned} \quad (6)$$

For the Reissner-Nordström metrics

$$g_{00} = 1 - \frac{2m}{r} + \frac{Q^2}{r^2} > 0 \quad \text{and} \quad g_{11} = -\frac{1}{g_{00}} < 0 \quad (7)$$

The matrix of the quadratic form (5) has the form

$$A = \begin{bmatrix} a_{00} & a_{01} & 0 & 0 \\ a_{01} & a_{11} & 0 & 0 \\ 0 & 0 & a_{22} & 0 \\ 0 & 0 & 0 & a_{33} \end{bmatrix} \quad (8)$$

The reduction of the quadratic form (5) to its canonical form is done through the method of eigenvalues and eigenvectors, because in this way we can determine an orthonormal basis  $B' = \{e'_0, e'_1, e'_2, e'_3\}$  corresponding to an isometry (rotation in  $\mathbb{R}^4$ ).

Solving the characteristic equation

$$p(\lambda) = (-1 - \lambda)^2 [\lambda^2 - (a_{00} + a_{11})\lambda + (a_{00}a_{11} - a_{01}^2)] = 0 \quad (9)$$

we obtain the eigenvalues

$$\lambda_0 = \frac{a_{00} + a_{11} + \sqrt{\Delta}}{2} > 0; \quad \lambda_1 = \frac{a_{00} + a_{11} - \sqrt{\Delta}}{2} < 0; \quad \lambda_2 = \lambda_3 = -1 \quad (10)$$

where

$$\Delta = (a_{00} - a_{11})^2 + 4a_{01}^2 > 0. \quad (11)$$

Therefore, metrics (5) becomes

$$ds^2 = \lambda_0 v_0^2 + \lambda_1 v_1^2 + \lambda_2 v_2^2 + \lambda_3 v_3^2. \quad (12)$$

After the reduction of metrics (5) to the canonical form (12) we will have the orthonormal basis

$$B' = \{e'_0, e'_1, e'_2, e'_3\} \quad (13)$$

which is obtained by norming the eigenvectors resulting from solving the characteristic system

$$\begin{cases} (a_{00} - \lambda)x_0 + a_{01}x_1 = 0 \\ a_{01}x_0 + (a_{11} - \lambda)x_1 = 0 \\ (-1 - \lambda)x_2 = 0 \\ (-1 - \lambda)x_3 = 0 \end{cases} \quad (14)$$

by replacing in turns the eigenvalues (10) in this system.

We get

$$\begin{cases} e'_0 = \left( \frac{1}{\sqrt{1+\alpha^2}}, \frac{-\alpha}{\sqrt{1+\alpha^2}}, 0, 0 \right) \\ e'_1 = \left( \frac{\alpha}{\sqrt{1+\alpha^2}}, \frac{1}{\sqrt{1+\alpha^2}}, 0, 0 \right) \\ e'_2 = (0, 0, 1, 0); \quad e'_3 = (0, 0, 0, 1) \end{cases} \quad (15)$$

$$B_0 = \{e_0 = (1, 0, 0, 0), e_1 = (0, 1, 0, 0), e_2 = (0, 0, 1, 0), e_3 = (0, 0, 0, 1)\} \quad (16)$$

to basis  $B^1$  is equal

$$C = \begin{bmatrix} \frac{1}{\sqrt{1+\alpha^2}} & \frac{+\alpha}{\sqrt{1+\alpha^2}} & 0 & 0 \\ \frac{-\alpha}{\sqrt{1+\alpha^2}} & \frac{1}{\sqrt{1+\alpha^2}} & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix} \quad (17)$$

and

$$\det C = 1 \quad (18)$$

where

$$\alpha = \frac{\sqrt{\Delta} - (a_{00} - a_{11})}{-2a_{01}} \quad (19)$$

From (18) it result that the present transformation is a rotation of the real space with four dimensions. The angular coordinates  $x^2$  and  $x^3$  are not modified.

The rotation angle,  $\gamma$ , is determined from the formulae:

$$\cos \gamma = \frac{1}{\sqrt{1+\alpha^2}}; \quad \sin \gamma = \frac{+\alpha}{\sqrt{1+\alpha^2}}; \quad \text{or } \text{tg} \gamma = +\alpha. \quad (20)$$

From relation (20), (19) and (11) it result that

$$\alpha = \text{tg} \gamma = \frac{\sqrt{(g_{00}^2 + 1 - w^2)^2 + 4w^2} - (g_{00}^2 + 1 - w^2)}{2w} \quad (21)$$

Taking into (21)  $m = 0, Q = 0$ , we get

$$\operatorname{tg} \gamma = \frac{\sqrt{w^4 + 4 + w^2} - 2}{2w} = \frac{w}{2} \left( 1 + \frac{1}{4}w^2 - \frac{1}{64}w^6 + \frac{1}{516}w^{10} \dots \right) \approx \frac{w}{2} \quad (22)$$

where in the first approximation is the half radial velocity in geometrized units:

$$w = \frac{V_{\text{rad}}}{c} \quad (23)$$

The Doppler's formula showst that

$$\frac{\Delta \lambda}{\lambda} = \frac{V_{\text{rad}}}{c} = w \quad (24)$$

and

$$\frac{\Delta \lambda}{\lambda} = 2 \operatorname{tg} \gamma \quad (25)$$

### References

1. SASS, I. H. A., 1986, Babeş-Bolyai Univ., Fac. Math. Res. Seminars, Preprint 6, Cluj-Napoca
2. LANG, K. R., "Astrofiziceskie formula"( Russian I.), Ed. Mir., Moskva

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